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# **Update on Simulations of Mercury Targets**

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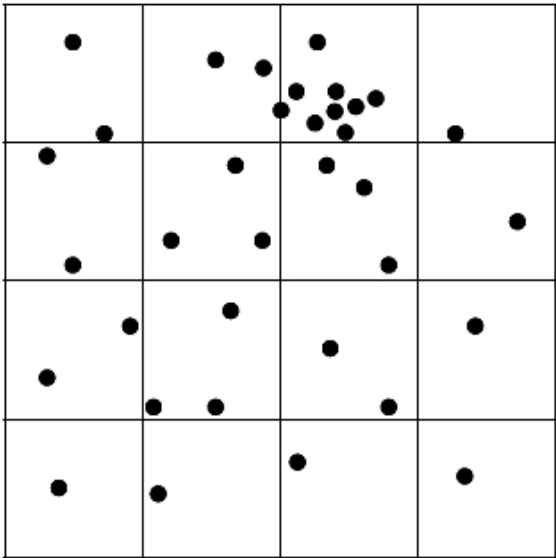
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Brookhaven National Laboratory*

# Outline

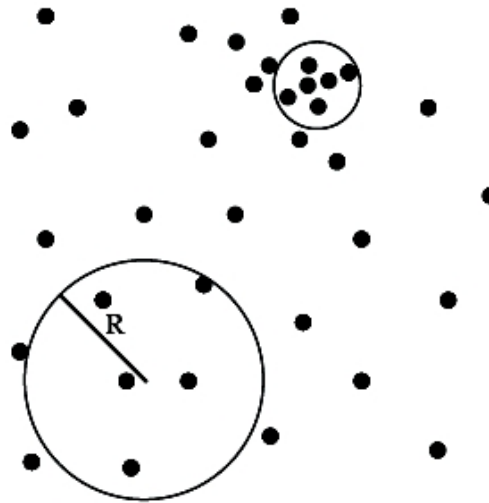
- SPH simulations of jets interacting with proton beams
- Fluid – structure interaction with SPH and mercury thimble simulations
- Front tracking simulations of free surface MHD at large density ratios

# Main Idea of SPH

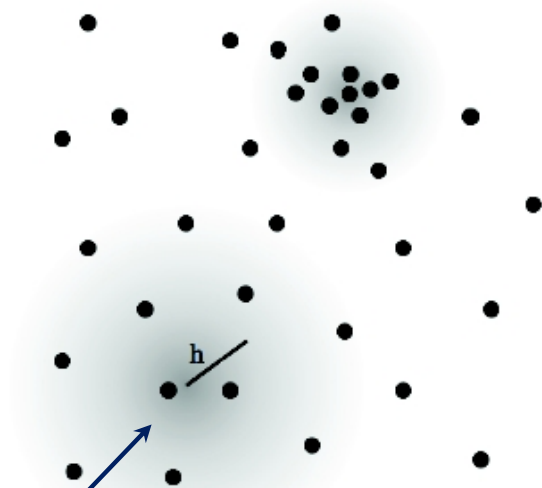
## Computing density of continuum using particles



Particle-mesh methods



Sum of particles in disks



- SPH:
- Density is weighted sum of particles
  - Each particle represents a Lagrangian cell
  - No particle connectivity

$$\rho_i = \sum_j m_j W_{ij}(h)$$

# Main Approach of SPH

- Kernel approximation: replace the delta-function with a smooth kernel function

$$A(\vec{r}) = \int A(\vec{r}') W(\vec{r} - \vec{r}', h) d\vec{r}'$$

- Approximate this integral using some particle distributions

$$A(\vec{r}) = \sum_b m_b \frac{A_b}{\rho_b} W_{ab}$$

- Discretize Navier-Stokes (or MHD) equations in Lagrangian form

$$\frac{D\vec{v}}{Dt} = -\frac{1}{\rho} \vec{\nabla} P + \vec{g} + \vec{\Theta} \quad \text{Momentum PDE in Lagrangian system}$$

Discretized Momentum Equation

$$\frac{d\vec{v}_a}{dt} = -\sum_b m_b \left( \frac{P_b}{\rho_b^2} + \frac{P_a}{\rho_a^2} + \Pi_{ab} \right) \vec{\nabla}_a W_{ab} + \vec{g}$$

# Benefits of SPH

- A parallel SPH hydro / MHD code has been developed
  - Collection of solvers, smooth kernels, EOS and other physics models
- Exact conservation of mass (Lagrangian code)
- Natural (continuously self-adjusting) adaptivity to density changes
- Capable of simulating extremely large non-uniform domains
- Ability to robustly handle material interfaces of any complexity
- Scalability on modern multicore supercomputers

# SPH Simulations

- Disruption of mercury targets interacting with proton pulses
- Entrance of spent mercury jets into the mercury pool

# Muon Collider vs Neutrino Factory

Beam: 8 GeV, 4 MW,  $3.125 \times 10^{15}$  particles/s, r.m.s. rad = 1.2 mm

Muon Collider:

15 bunches / s

66.7 ms interval

208 teraproton per bunch

Neutrino Factory:

150 bunches / s

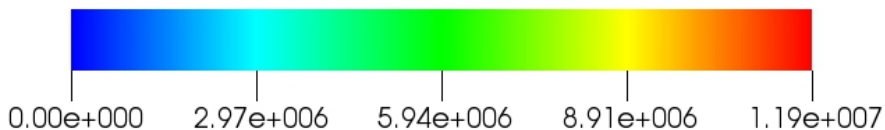
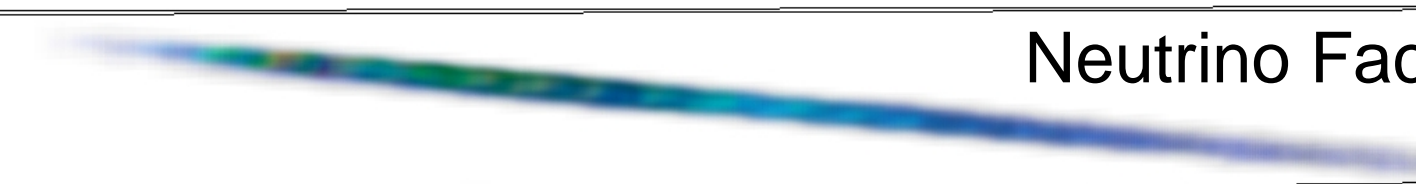
6.67 ms interval

20.8 teraproton per bunch

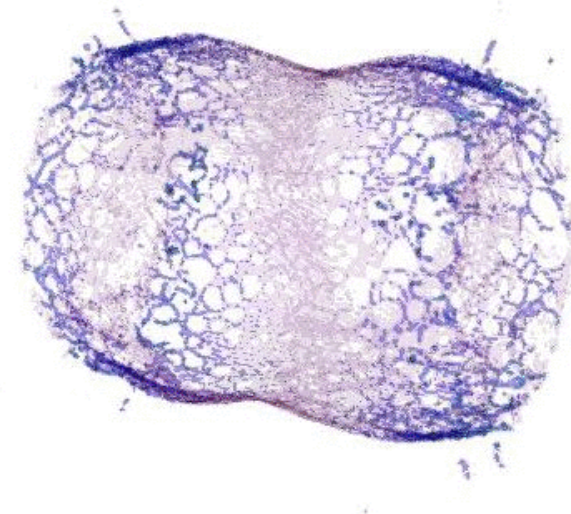
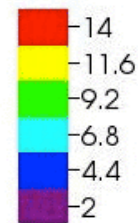
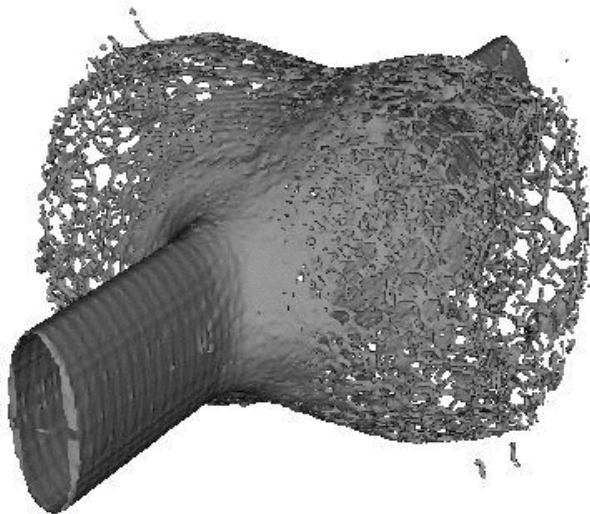
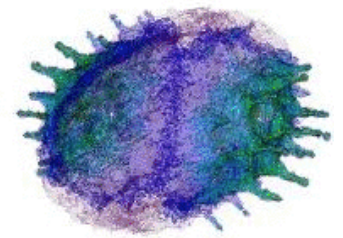
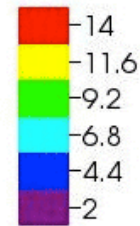
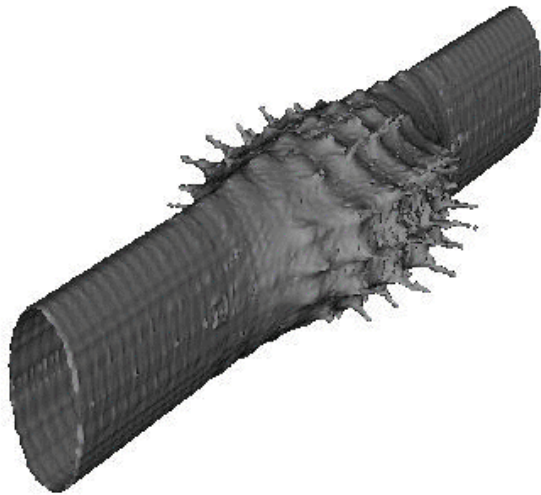
Maximum pressure (estimate):

Muon Collider:  $P_{\max} = 110$  kbar

Neutrino Factory:  $P_{\max} = 11$  kbar



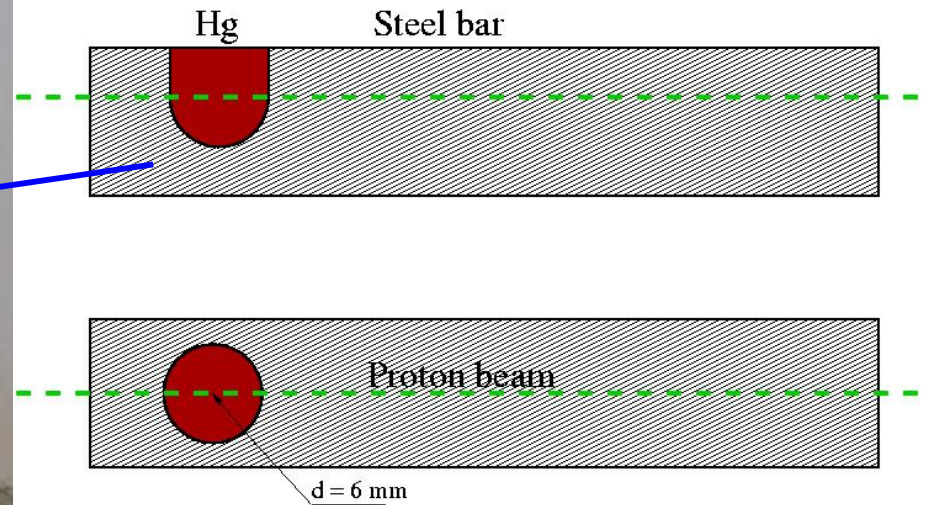
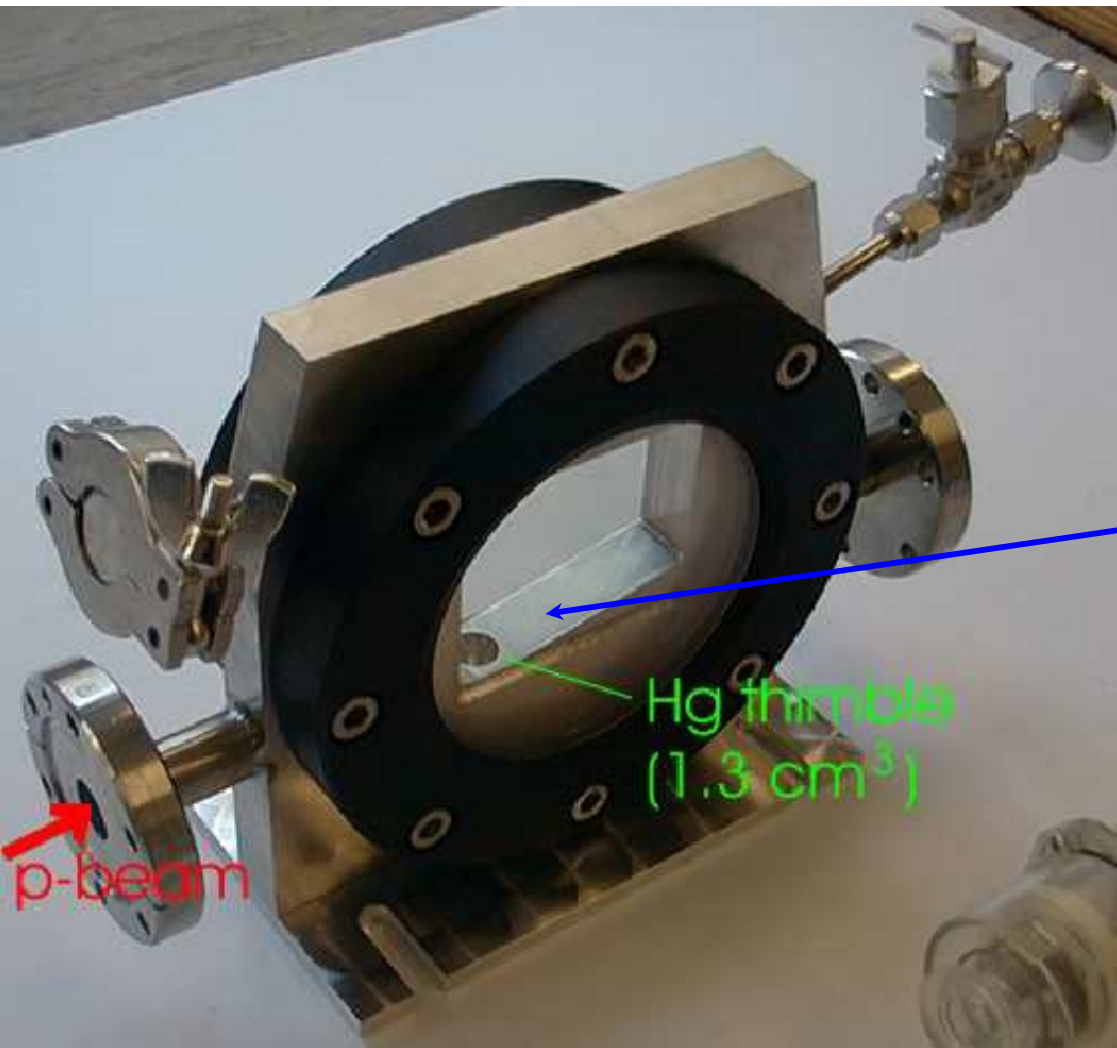
# Mercury Jet after Interaction with Proton Pulse



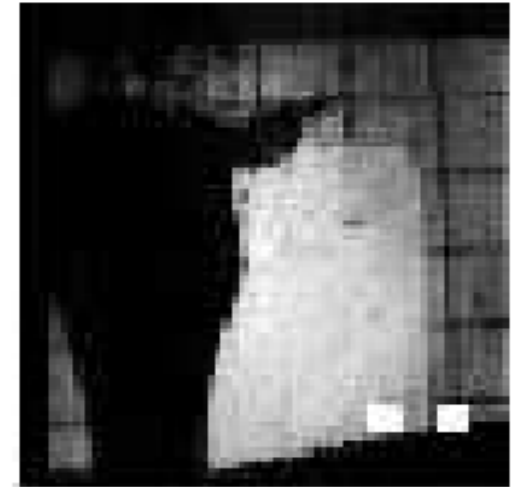
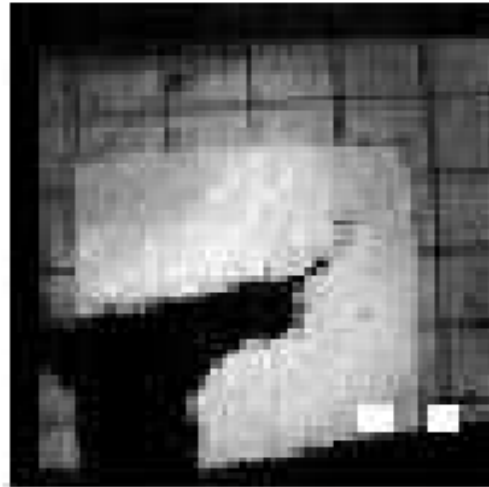
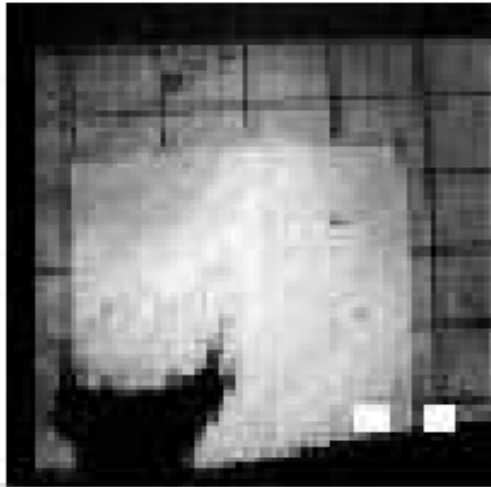


# **SPH Simulations of mercury thimble experiments**

# Experimental Setup

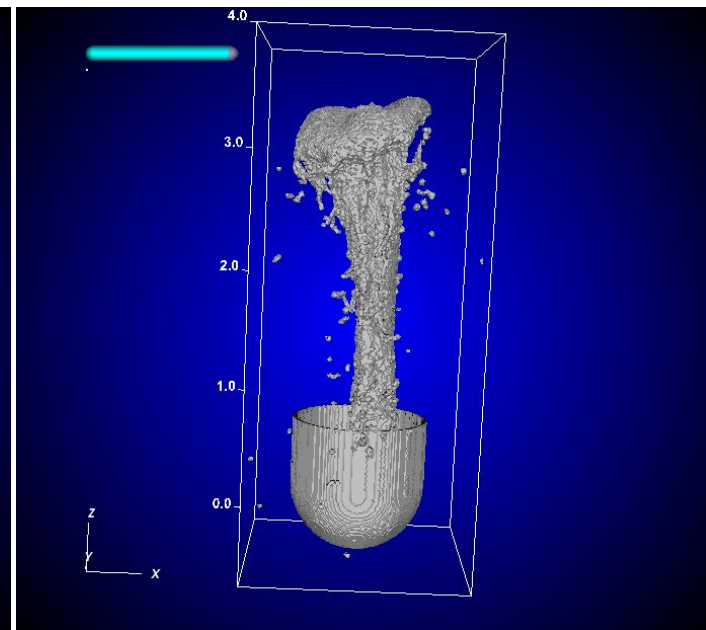
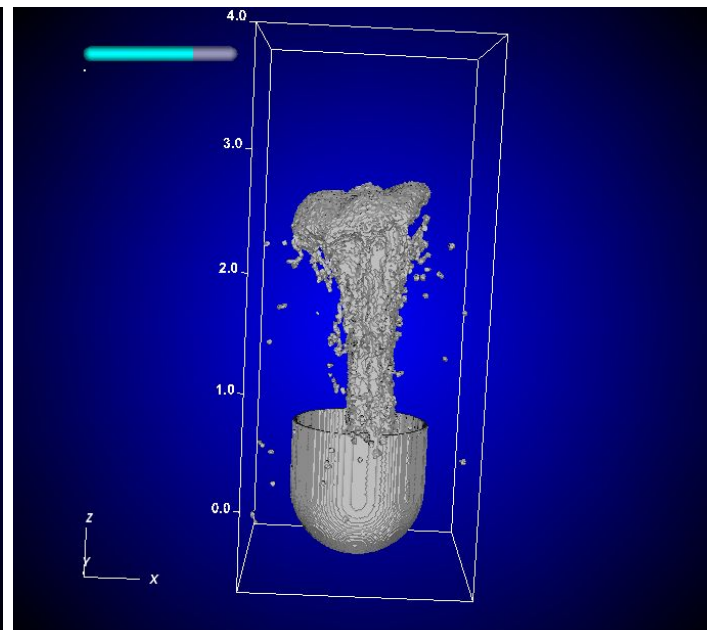
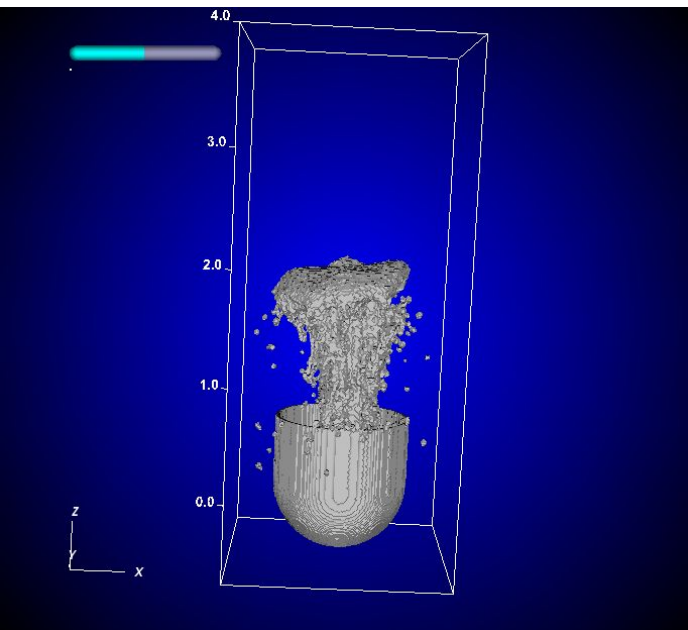
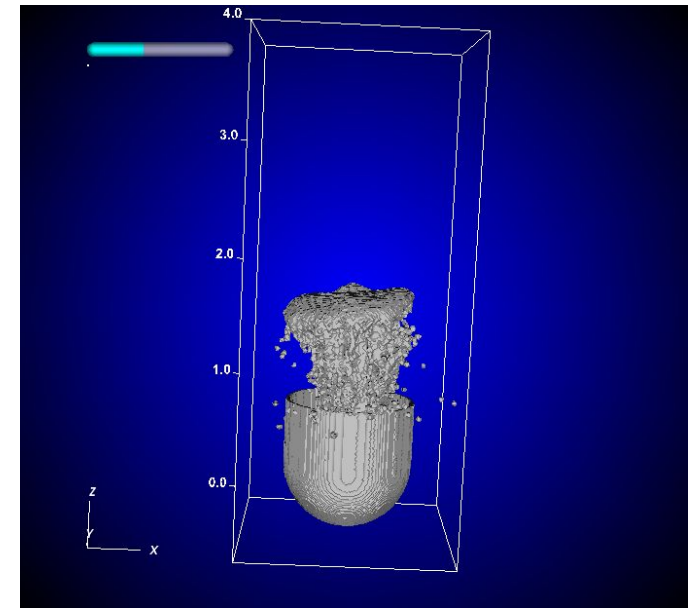
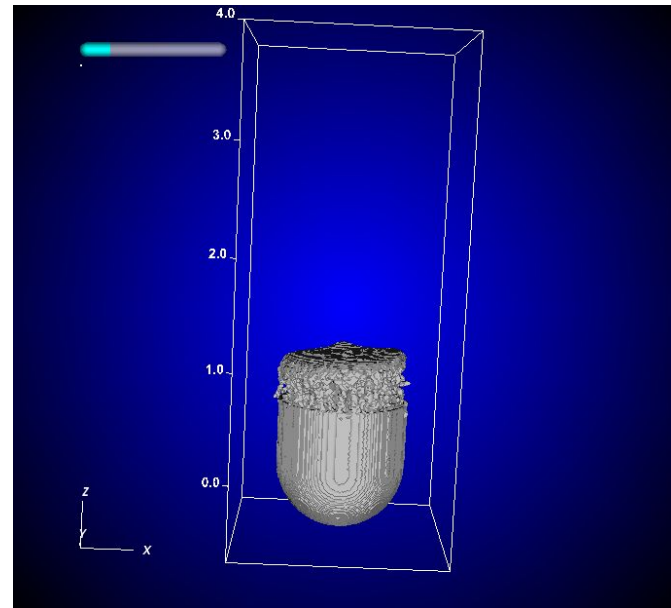
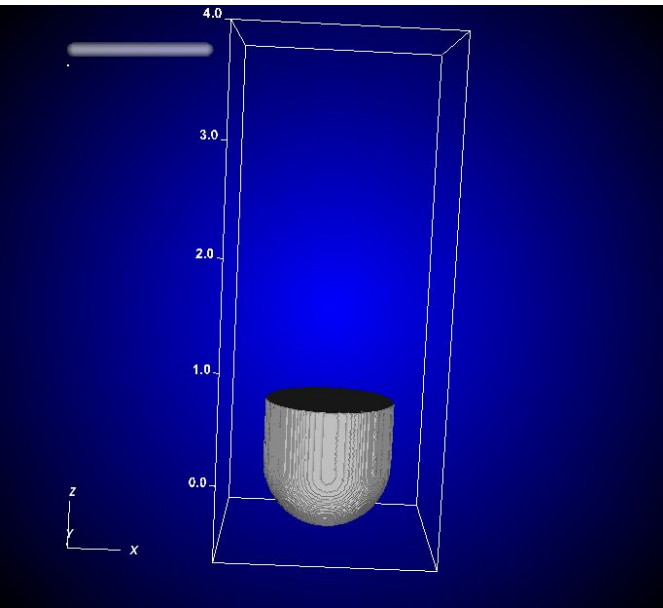


# Typical Experimental Results

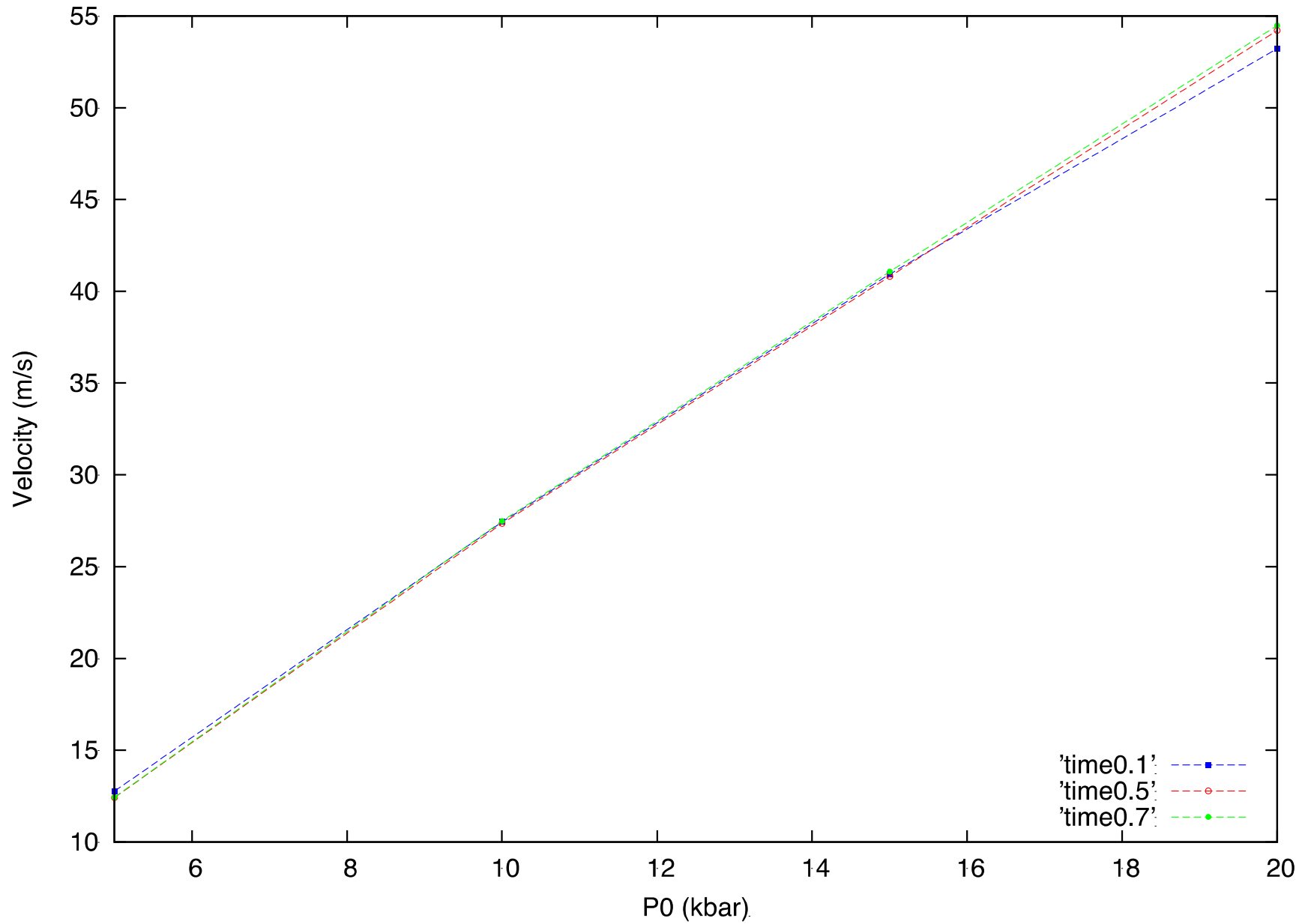


Mercury splash at  $t = 0.88, 1.25$  and  $7$  ms after proton impact of  $3.7$  teraprotons

# Simulation results



# Simulation results



# **FronTier Simulations of Incompressible MHD at Large Density Ratios**

# Equations of Incompressible MHD at Low Magnetic Reynolds Numbers

$$\rho \left( \frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla \right) \mathbf{u} = \mu \Delta \mathbf{u} - \nabla P + \rho \mathbf{g} + \frac{1}{c} (\mathbf{J} \times \mathbf{B})$$

$$\nabla \cdot \mathbf{u} = 0$$

$$\mathbf{J} = \sigma \left( -\nabla \varphi + \frac{1}{c} \mathbf{u} \times \mathbf{B} \right)$$

$$\nabla \cdot \mathbf{J} = 0$$

$$\nabla \cdot (\nabla \varphi) = \frac{1}{c} \nabla \cdot (\mathbf{u} \times \mathbf{B})$$

$$\left. \frac{\partial \varphi}{\partial \mathbf{n}} \right|_{\Gamma} = \left. \frac{1}{c} (\mathbf{u} \times \mathbf{B}) \cdot \mathbf{n} \right|_{\Gamma}$$

# Equations of Incompressible MHD at Low Magnetic Reynolds Numbers

$$\frac{\partial \rho}{\partial t} = -\nabla \cdot (\rho \mathbf{u})$$

$$\rho \left( \frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla \right) \mathbf{u} = -\nabla P + \mu \Delta \mathbf{u} + \frac{1}{c} (\mathbf{J} \times \mathbf{B})$$

$$\rho \left( \frac{\partial}{\partial t} + \mathbf{u} \cdot \nabla \right) e = -P \nabla \cdot \mathbf{u} + \frac{1}{\sigma} \mathbf{J}^2$$

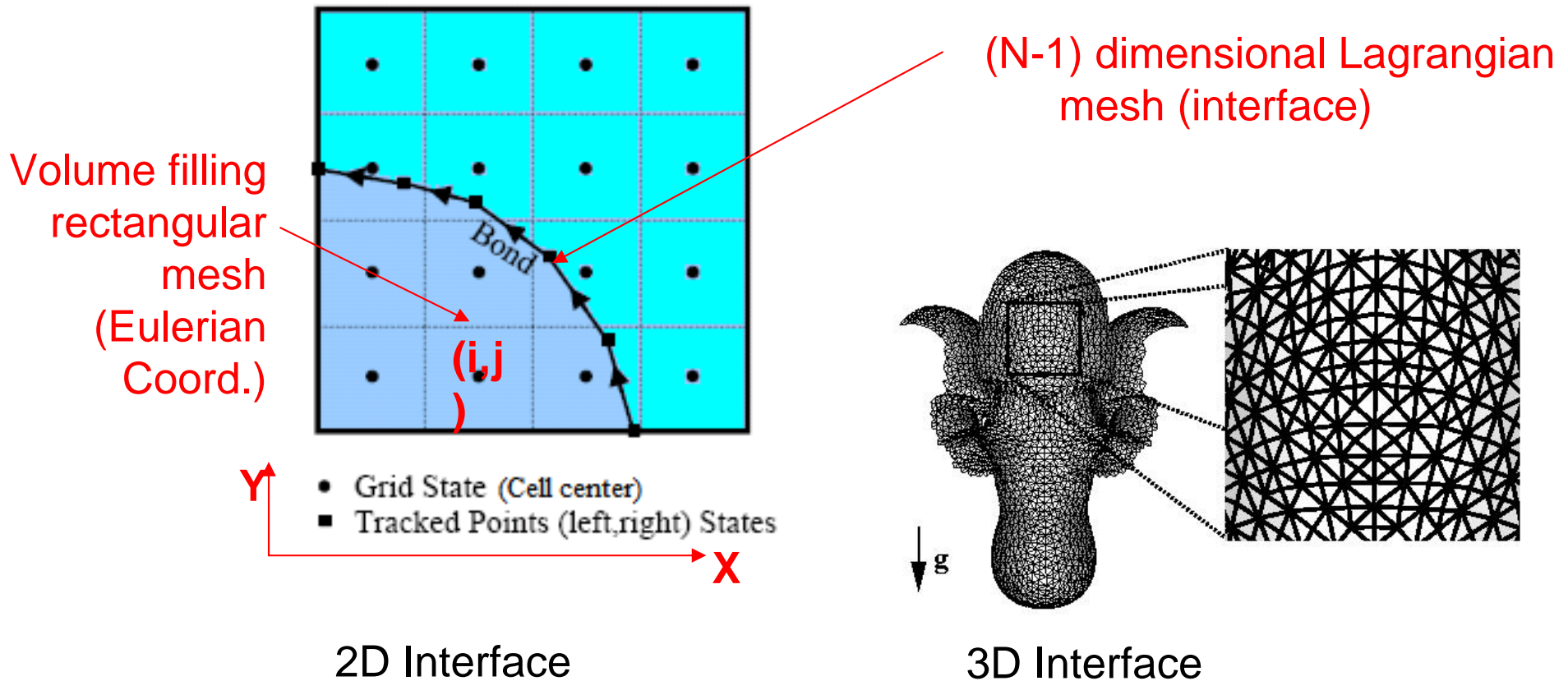
$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{u} \times \mathbf{B}) - \nabla \times \left( \frac{c^2}{4\pi\sigma} \nabla \times \mathbf{B} \right)$$

$$P = P(\rho, e), \quad \nabla \cdot \mathbf{B} = 0$$



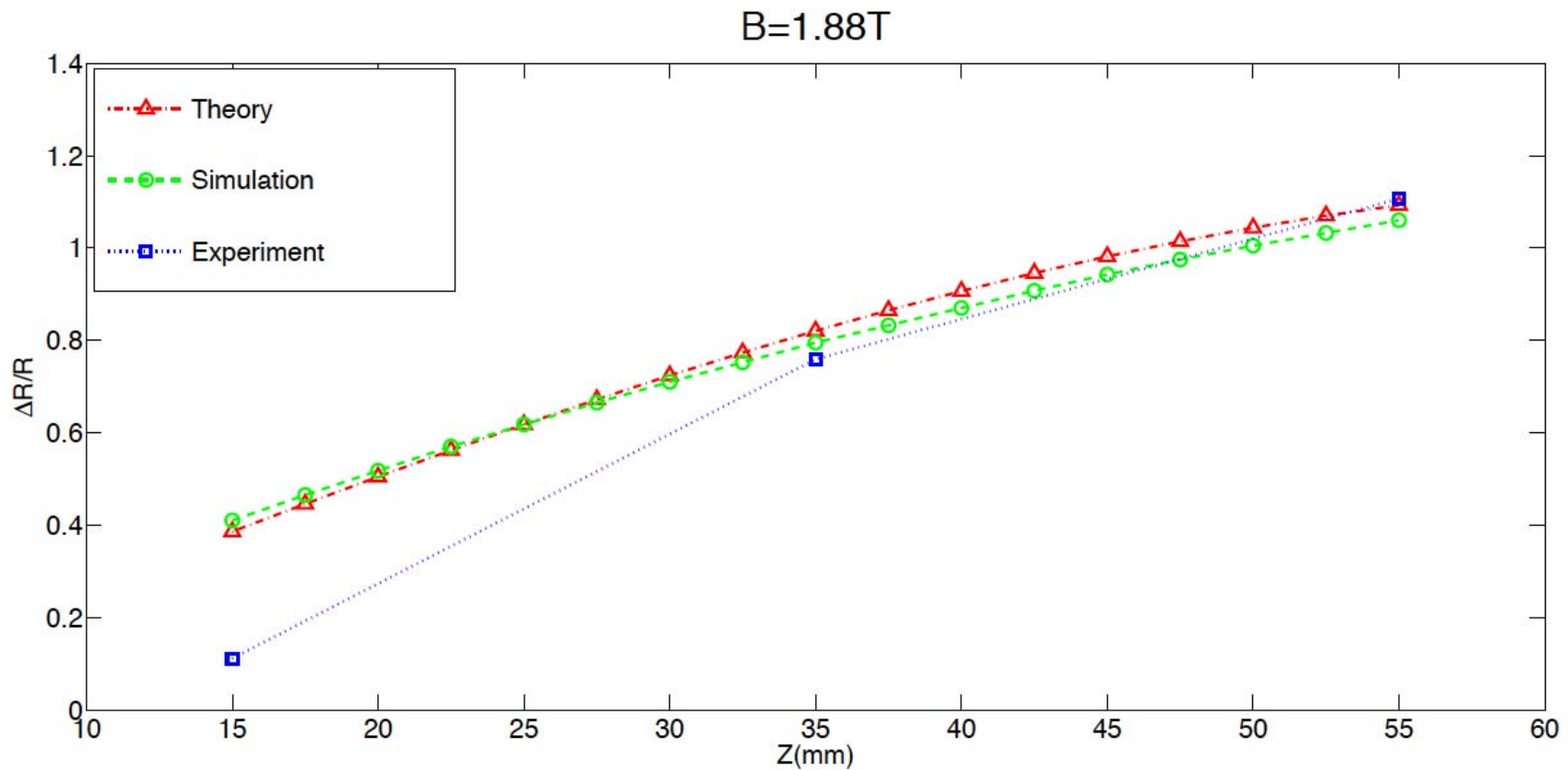
# Main Idea of Front Tracking

- Front tracking is a hybrid Lagrangian-Eulerian method for systems with sharp discontinuities in solutions or material properties



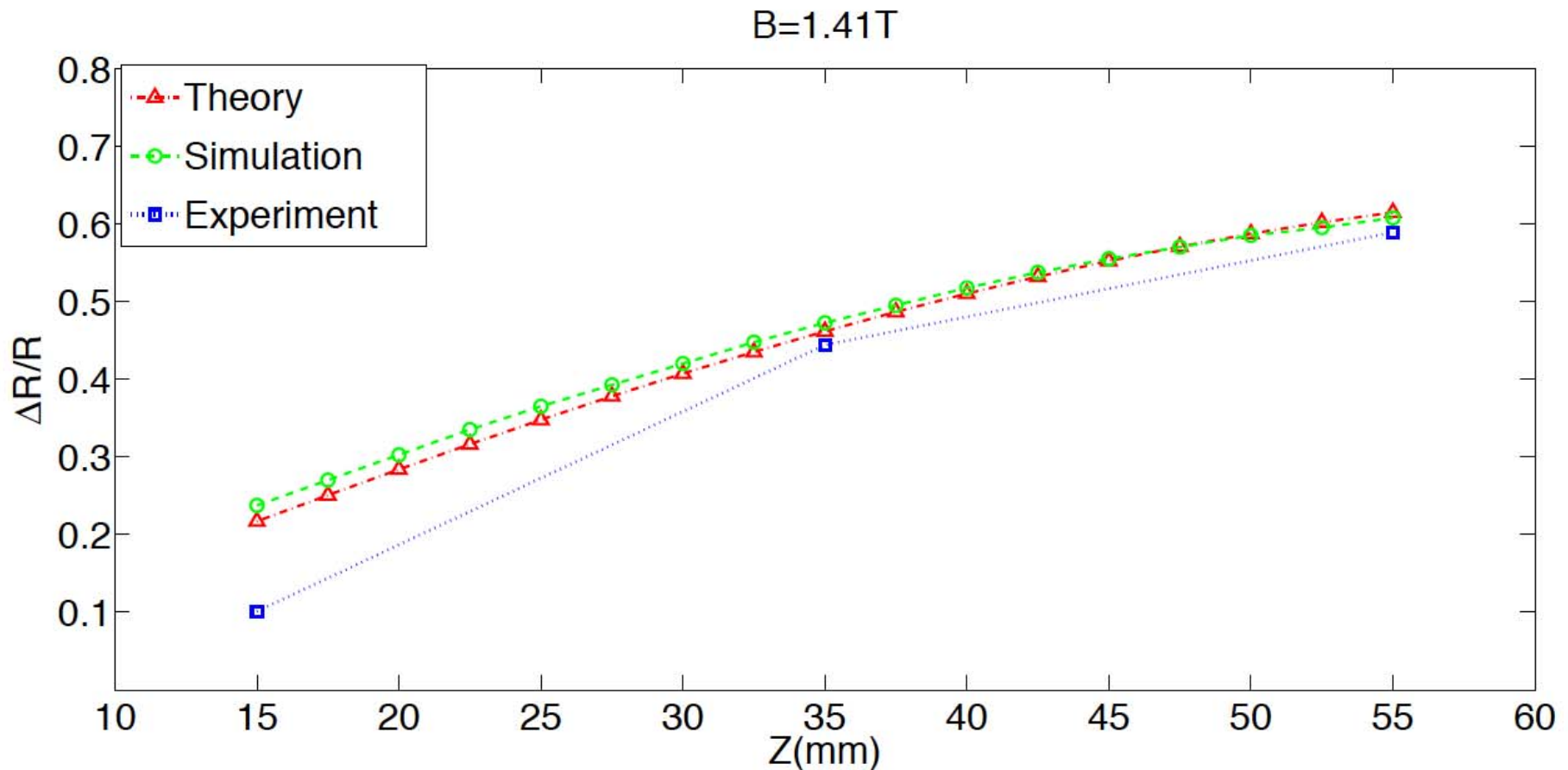
# Verification and Validation: Mercury Jet in Transverse Magnetic Field

## Real density ratio, sharp interface



# Verification and Validation: Mercury Jet in Transverse Magnetic Field

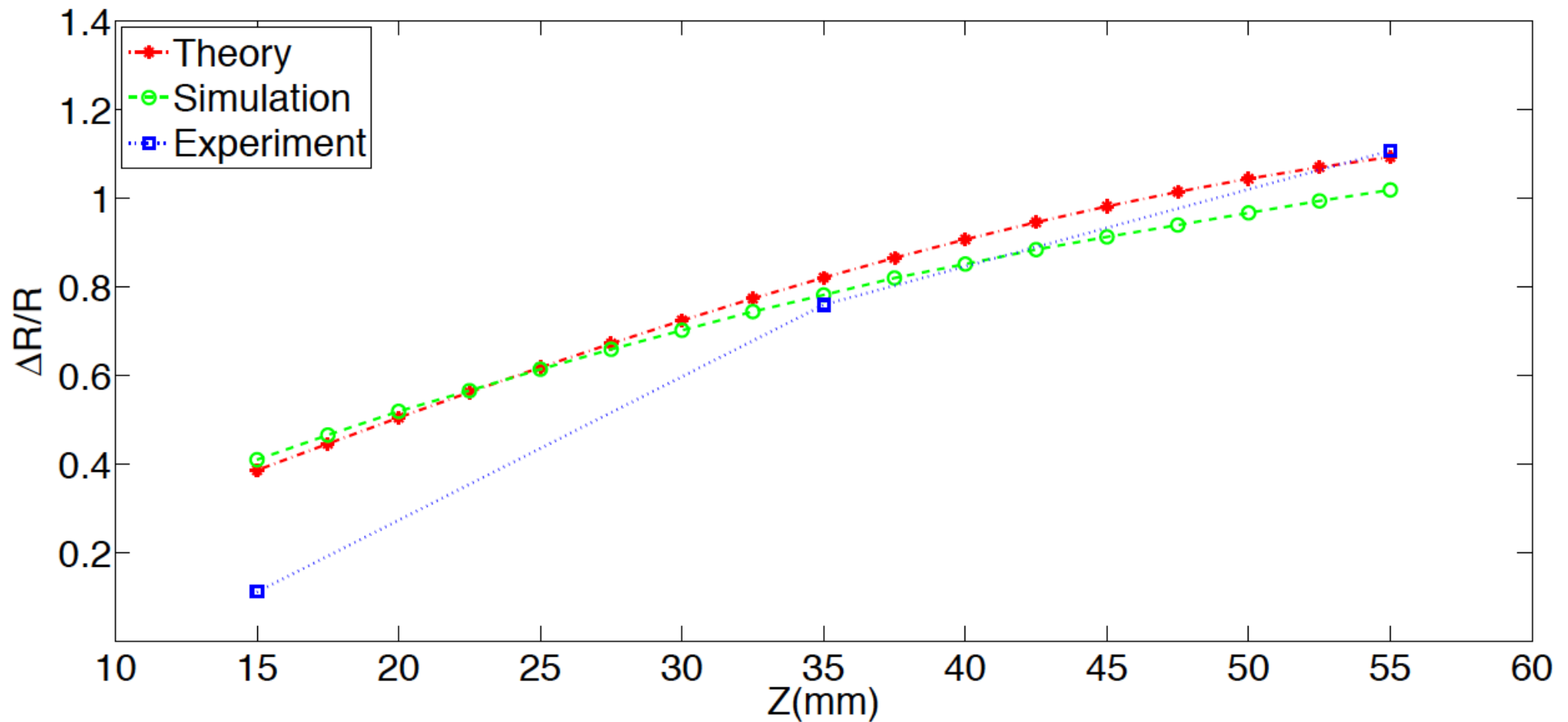
## Real density ratio, sharp interface



# Verification and Validation: Mercury Jet in Transverse Magnetic Field

## Density ratio 10, smoothed interface

B=1.88T



# Verification and Validation: Mercury Jet in Transverse Magnetic Field

## Density ratio 10, smoothed interface

