

6/4/82

DEAR KIRK

FROM NATHAN MYHRVOLD

I'M SORRY I HAVEN'T BEEN QUICKER ABOUT FINISHING UP THOSE CALCULATIONS, BUT I'M SUPPOSED TO FINISH UP ~~THE~~ PAPER IN TIME FOR MALCOLM'S RETURN (SOME TIME THIS WEEK OR NEXT). I'LL TRY HAVE SOMETHING MORE DETAILED THAN THIS NOTE FOR YOU BY WEDNESDAY OR THURSDAY.

I'VE BEEN LOOKING INTO SEVERAL THINGS:

1. CIRCULAR MOTION

ALTHOUGH THE POWER STILL LOOKS TOO LOW IN THE PEAK RANGE OF THE SPECTRUM, THE TAIL (HIGH ENERGY) END OF THE SPECTRUM LOOKS PROMISING. IT IS MUCH FASTER THAN ~~THE~~ FOR A THERMAL SPECTRUM AT THE SAME PEAK ENERGY. THE PAPER BY LUTAN & PFAUSTICA, WHICH YOU MENTIONED HAS GRAPHS OF THE SPECTRUM (LOOK AT THE "CONSTANT ACCELERATION" GRAPH). ALTHOUGH L & P BLEW IT ON THE REST OF THEIR PAPER, THEIR GRAPH IS O.K..

ANYWAY, THERE IS HOPE THAT AT SUFFICIENTLY HIGH ENERGY THE RADIATION FROM OUR PROCESS WILL BECOME COMPARABLE TO SYNCHROTRON RADIATION - I'LL WORK ON CHECKING THIS OUT MORE.

2. LINEAR MOTION - I'M GOING TO WRITE UP A LITTLE CALCULATION TO SEE IF WHAT WE CAN EXPECT FROM THIS CASE.

3. Quantum Mechanical Approach

Your question about the full quantum mechanical approach to the "acceleration radiation" is both interesting and difficult, for quite some time I've been wanting to work out a Q.M. version of this stuff, but I haven't found the time. I'll try to do some preliminary work ~~this~~ ~~work~~ ~~etc.~~

V. A. K.

KIRK:

6/23/92

THE CORRECT ACCUMULATION a TO USE
IN $T = \frac{h a}{2\pi k c}$ DOES SEEM TO BE

THE ACCUMULATION AS MEASURED IN THE
REST FRAME OF THE ACCUMULATING
OBJECT. IF THE ACCUMULATION IS
COLLINEAR WITH THE VELOCITY THEN IT
IS RELATED TO THE LAB ACC. BY γ^3
AS YOU SAID. ~~W/~~

HOWEVER ~~+~~ SINCE THE TEMPERATURE AND INTENSITY
OF THE "RADIATION" YOU GET FROM

$$T = \frac{h a}{2\pi k c}$$

IS IN THE REST FRAME OF THE OBJECT,
IT MUST THEREFORE

~~BE COMPARED~~ BE COMPARED TO OTHER
INCOMING STUFF WHICH IS BLUE SHIFTED
BY A FACTOR γ .

I'D LIKE TO TALK TO TALK TO
YOU ABOUT THIS SOME TIME SOON.

I'M USUALLY IN IN THE ^{MID TO} LATE AFTERNOONS (1:30 TO 5:00)
(SOME TIMES I'M IN THE STUDENT SHOP
WORKING ON MY CANNON), I ALSO AM
HERE LATE AT NIGHT (9 PM TO 4 OR 6 AM).

DROP A NOTE IN MY BOX IF YOU WANT
TO ARRANGE SOMETHING - I AM WILLING
TO COME IN MORNINGS IF ITS REALLY
NECESSARY.

I MADE COPIES OF THE TWO
SCHEMA PRINTS FOR YOU. ALTHOUGH
SOME OF THE TREATMENT ~~IS~~ ISN'T VERY
RELEVANT UNLESS YOU'RE USED TO A
PARTICULAR SET OF JARGON (SUCH AS QUANTUM
OPTICS) - THEY DO GIVE A GOOD INTRODUCTION

TO THE SUBNET.

A handwritten signature in dark ink, appearing to be 'N. H. ...', written in a cursive style.

DEAR KIRK

Here is the first installment
of my calculations. Unfortunately
I had severe Xerox machine problems
so this is only part of what I have.

I thought it would be useful
to write down a ~~few~~ basic summary
of the theoretical approaches + references.
I have also included some of my
speculations on ~~how to go~~ the
ways to distinguish the thermal
radiation from ^{the} synchrotron or bremsstrahlung
background. I'll have my numerical
estimates for you soon.

I've found a pretty good
book on synchrotron radiation:
SYNCHROTRON RADIATION BY SOKOLOV
AND TERNOV 82/6. 655. 1968.

The library has several copies if
you want to look at one.



Princeton University

DEPARTMENT OF PHYSICS: JOSEPH HENRY LABORATORIES

JADWIN HALL

POST OFFICE BOX 708

PRINCETON, NEW JERSEY 08544

July 8, 1983

Prof. J. Schwinger
Dept. of Physics
U.C.L.A.
Los Angeles, CA 90024

Dear Prof. Schwinger,

Some recent results of general relativity have caused renewed interest in your work of the early 50's on electrons in strong electromagnetic field. The new issues are intriguing but rather technical and I would appreciate any advice you would care to give.

The basic idea of Fulling, Davies, and Unruh is that the Hawking radiation effect from black holes has an equivalent effect for accelerated observers in zero gravitational field. Further, a charged, accelerating observer such as an electron can scatter off the thermal bath of radiation leading to light detectable by an inertial observer. Of course this 'scattering' process as viewed by an inertial observer could also be considered as a correction to more ordinary forms of radiation of an accelerating charge. A semiclassical model uses the Hawking temperature

$$T = \frac{\hbar a}{2\pi c k}$$

(a = rest frame acceleration)

to suggest an electron would radiate an additional power of

$$\frac{dU}{dt} = \frac{\hbar \nu_0^2 a^4}{90\pi c^6} \quad \left(\nu_0 = \frac{e^2}{m c^2} \right)$$

The dependence of the radiation rate on the fourth power of the acceleration is indicative of an underlying thermal process.

For the new effect to be sizable the acceleration must be very large indeed, $>10^{31} g$. Aside from elementary particle collisions in which the concept of acceleration is doubtful, such large accelerations might be achieved in 'collisions' of the SLAC electron beam with a very high intensity laser beam.

A recent thesis by Nathan Myhrvold at Princeton examines the propagation of an electron in a strong static electric field and concludes the thermal radiation effect is indeed present, based on your earlier work. However the study was not carried far enough to indicate how the thermal effects might be clearly distinguished from other QED processes.

Any comment or advice you would care to give on this matter would be most welcome.

Sincerely yours,

Kirk McDonald
Assoc. Prof. of Physics

From: SMTP% "unruh@physics.ubc.ca" 23-APR-1998 13:24:24.03
To: "Kirk T. McDonald 609-258-6608" <mcdonald@puphed.princeton.edu>
CC:
Subj: Re: My paper

Return-Path: unruh@physics.ubc.ca
Received: by puphed.princeton.edu (UCX V4.1-12, OpenVMS V7.0 VAX);
Thu, 23 Apr 1998 13:24:22 -0400
Received: from theory.physics.ubc.ca (black-hole.physics.ubc.ca [137.82.43.40])
by feynman.princeton.edu (8.8.8/8.8.8) with ESMTP id NAA26069
for <mcdonald@puphed.princeton.edu>; Thu, 23 Apr 1998 13:24:23 -0400 (EDT)
Received: from localhost (unruh@localhost) by theory.physics.ubc.ca (8.8.5/8.8.3) with SMTP id KAA02779 for
<mcdonald@puphed.princeton.edu>; Thu, 23 Apr 1998 10:24:22 -0700 (PDT)
X-Authentication-Warning: black-hole: unruh owned process doing -bs
Date: Thu, 23 Apr 1998 10:24:22 -0700 (PDT)
From: Bill Unruh <unruh@physics.ubc.ca>
X-Sender: unruh@black_hole
To: "Kirk T. McDonald 609-258-6608" <mcdonald@puphed.princeton.edu>
Subject: Re: My paper
In-Reply-To: <98042312565836@puphed.princeton.edu>
Message-ID: <Pine.SUN.3.95.980423101835.2775A-100000@black_hole>
MIME-Version: 1.0
Content-Type: TEXT/PLAIN; charset=US-ASCII

Yes, it turned out to be somewhat more complicated than I expected!. On the other hand I at least feel that it gives a much better physical insight into what is happening than does the classical approach. In particular, I have had a long argument with Jackson as to whether or not the classical approach requires that the particle move in the z direction-- Jackson claiming that it does not. (Bell and Leinaas and my approach make clear that that z motion is absolutely essential for explaining the depolarisation details) I guess that is my main problem with the classical approach- it is simply a calculation which produces almost no insight into what is going on, at least for me (Me? Biased?).

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On Thu, 23 Apr 1998, Kirk T. McDonald 609-258-6608 wrote:

> Received: from physics.ubc.ca (physics.ubc.ca [137.82.43.9]) by theory.physics.ubc.ca (8.8.5/8.8.3) with ESMTP id JAA02734 for <unruh@black-hole.physics.ubc.ca>; Thu, 23 Apr 1998 09:59:26 -0700 (PDT)
> Received: from feynman.princeton.edu (root@feynman.Princeton.EDU [128.112.100.5])
> by physics.ubc.ca (8.8.5/8.8.5) with ESMTP id JAA25676
> for <unruh@physics.ubc.ca>; Thu, 23 Apr 1998 09:59:17 -0700 (PDT)
> Received: from PUPHED (puphed.hep.princeton.edu [128.112.100.25])
> by feynman.princeton.edu (8.8.8/8.8.8) with SMTP id MAA23968
> for <unruh@physics.ubc.ca>; Thu, 23 Apr 1998 12:59:14 -0400 (EDT)
> Date: Thu, 23 Apr 1998 12:56:58 -0400
> Message-ID: <98042312565836@puphed.princeton.edu>
> From: "Kirk T. McDonald 609-258-6608" <mcdonald@puphed.princeton.edu>

> To: unruh@physics.ubc.ca
> Subject: Re: My paper
> X-VMS-To: SMTP%"unruh@physics.ubc.ca"
>
> Bill,
>
> I did pick your paper off the Web and have read it with interest --
> tho without mastery of the details.
>
> It looks good! But one might be pressed to say that it's simpler than
> the `classic' approach.
>
> Since the Debate with Jackson was about the validity of the approach, not
> the complexity, I'd say the Hawking-Unruh view is holding its own....
>
> --Kirk
>

From: SMTP%"finnr@fermi.phys.washington.edu" 5-MAR-1999 14:07:55.85
To: "Kirk T. McDonald 609-258-6608" <mcdonald@puphed.princeton.edu>
CC:
Subj: Re: strong fields

Kirk:

My Klein's Paradox paper is hidden away in Physica Scripta 23, 1030 (1981). A similar paper by Barry Holstein can be found in Am.J.Phys. 66, 507 (1998). The Leinaas of the Bell-Leinaas paper on the Unruh effect for circularly accelerated particles is a colleague of mine in Oslo.

I've looked at the papers in your accel subdirectory. Which one in particular is the version you wrote for Am.J.Phys. and where you had problems with the referees?

Finn R.

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